Reconstruction of the heliospheric magnetic field utilizing an accelerated computational grid

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Overview

- (accelerated) computational grid
- Mathematical description of the magnetic field (Potential Field Source Surface model)
- What has been done so far?
- Utilizing the results

(Accelerated) computational grid

 Finite computational and storage capacity demands computations take place on a finite number of grid points distributed over the domain of interest (here: near sun heliosphere) → methods of Computational Fluid Dynamics (CFD) and Magnetohydrodynamics (MHD)

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 Physics "happen" at the grid points, values in between assumed to be interpolated linearly (if that is not good enough, increase number of grid points)

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phi \rightarrow 72 – 360

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Mathematical description of the magnetic field

 As a first approximation and proof-of-concept for the computational grid, we are using a potential field source surface (PFSS) model

• If the need arises, more complicated equations can be incorporated into the grid (CSSS, full MHD-approach, ... ?)

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Disadvantages:

- Crude model (imposed geometrical symmetry, no dynamic processes for an entire Carrington rotation,...)
- Current free approach is only valid for the quiet sun (at most)

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- Mapping of magnetic configurations in 3D (Thies Peleikis will tell you more)

Utilizing the results

... or what is that data useful for?

- Mapping heliospheric in-situ measurements to corresponding photospheric locations (again: Thies Peleikis will tell you more)
- As a seed for solar wind models

Utilizing the results



so..... let's take a look!

Thank you for attention.

Questions? Suggestions?

Hairy Ball - WSO



Hairy Ball - KPVT



Hairy Ball – WSO with ballistic mapping



Hairy Ball – KPVT with ballistic mapping



Ballistic backmapping - WSO



Ballistic backmapping - KPVT



Ballistic backmapping - KPVT

